

**Working with Broadband Transient Absorption Data in Igor Pro
v. 3.2 March 9, 2011**

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Please direct questions and comments to the author:

Matt Sfeir

msfeir@bnl.gov

Starting Igor Pro

Version Requirements

These procedures require Igor Pro version 6.2 or later. Igor Pro is written by

Wavemetrics, Inc. More information can be found at:

<http://www.wavemetrics.com/>

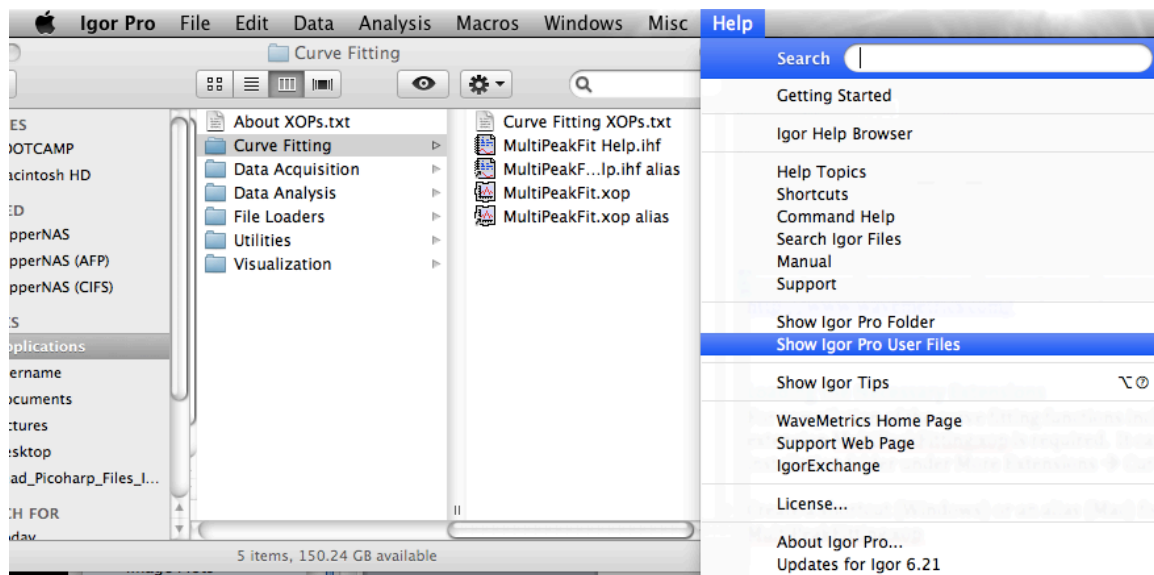
Loading the Necessary Extensions

For compilation of the curve fitting functions included in the procedure, the optional extension MultiPeakFit.xop is required. It can be found in the Igor Pro installation folder under More Extensions → Curve Fitting → MultiPeakFitting.xop

Create a shortcut (Windows) or an alias (Mac) for the files:

MultiPeakFit.xop

MultiPeakFit Help.ihf



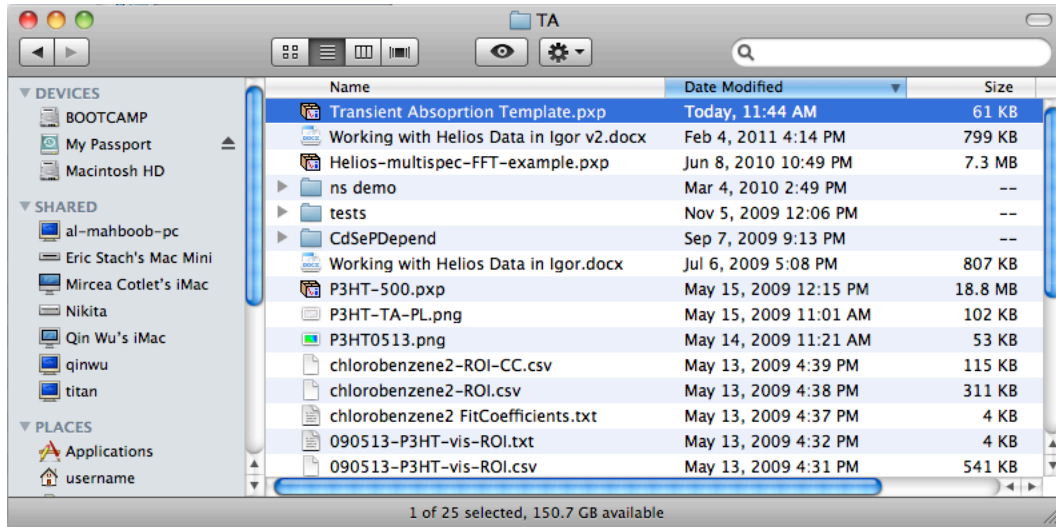
From the Igor Help menu, choose, “Show Igor Pro User Files.” Add these shortcuts to the Extensions folder in the User files directory.

If Igor is already running, then it must be restarted for these changes to take effect.

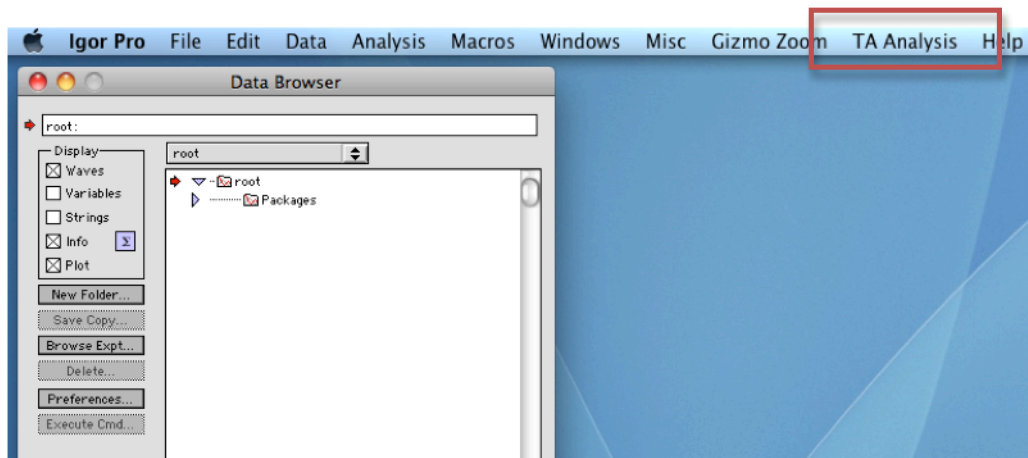
Importing Transient Absorption Data into Igor Pro

Open File

- 1) Open the Igor Packed Experiment file named “Transient Absorption Template.pxp”

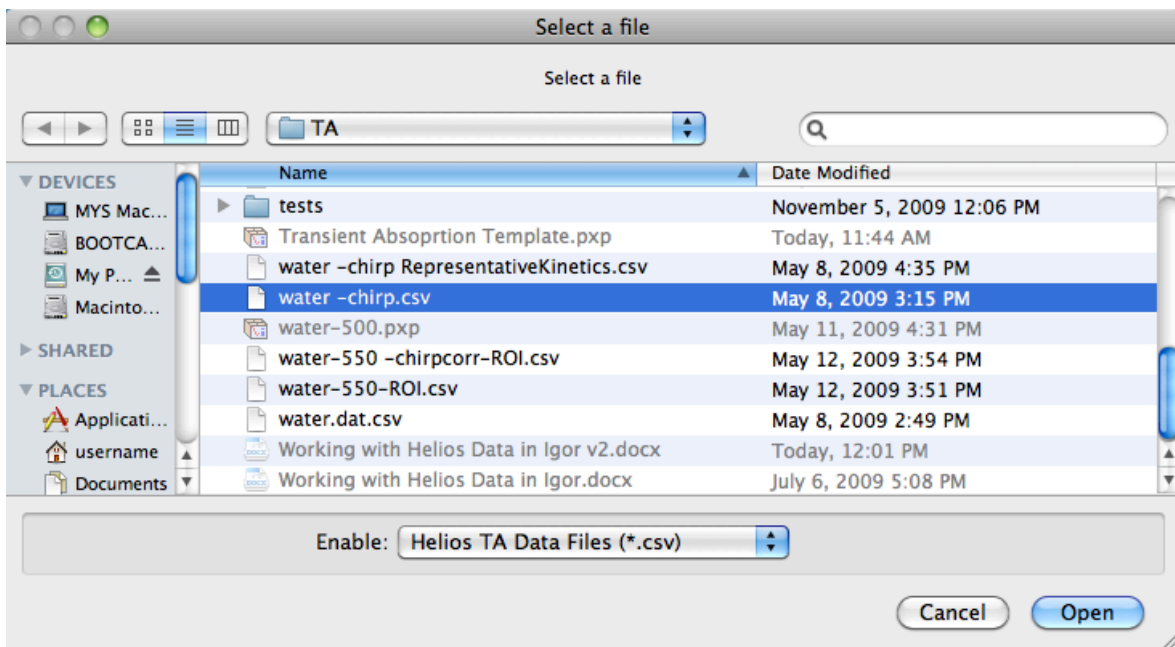
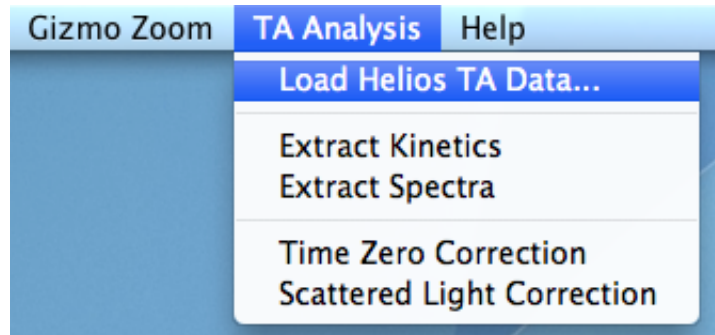


- 2) A blank experiment will open which automatically includes and compiles the necessary procedures. A new menu will appear on the task bar entitled, “TA Analysis.”



- 3) Open your saved transient absorption data (.csv files) using the appropriate menu option:

TA Analysis → Load Helios TA Data ...



Wave Format

- 4) This procedure loads the original data in *filename* with the corresponding wavelength (RP_*filename*) and time (CP_*filename*) information. It creates a new set of these waves (*filename_zmat*; *filename_xc*; *filename_yc*), with all the NaNs and Infs in the files replaced with 0, and deletes extra rows (> 512). It also creates several additional waves for graphing puposes:

RP_*filename* (wavelength labels for row data in nm)

filename (original uncorrected matrix in ΔOD)

CP_*filename* (time labels for column data in ps)

filename_zmat (corrected matrix)

filename_xc (processed row values)

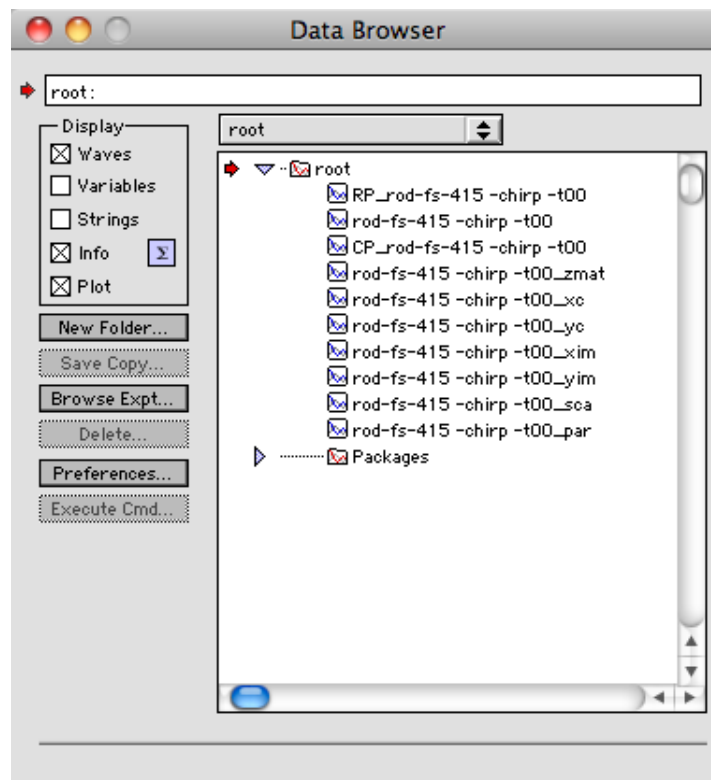
filename_yc (processed column values)

filename_xim (x values for image plot)

filename_yim (y values for image plot)

filename_sca (xyz wave for scatter plot)

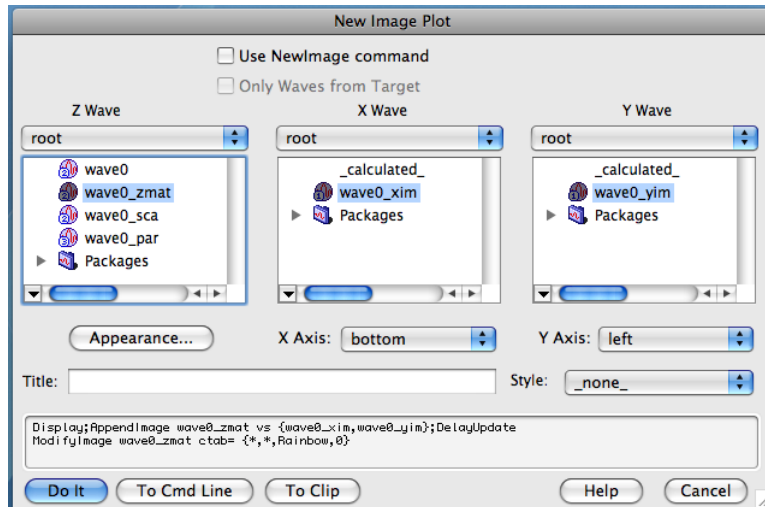
filename_par (3D parametric wave)



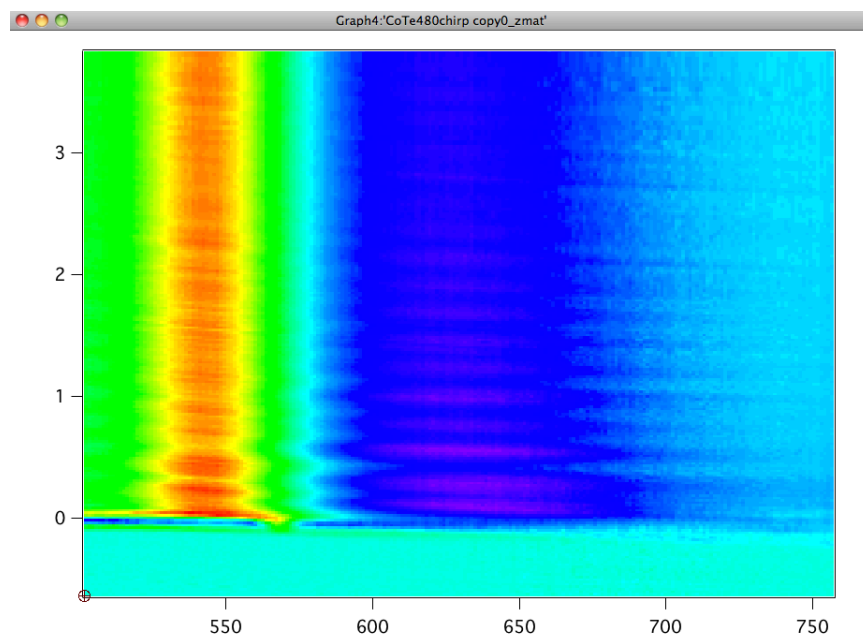
Plotting TA Data

Automatically Generated Plots

- 1) An image plot of your data will be automatically created with a visible cursor.
This plot can be generated manually in the following way, using *filename_zmat*, *filename_xim*, and *filename_yim*:
Windows → New → Image Plot



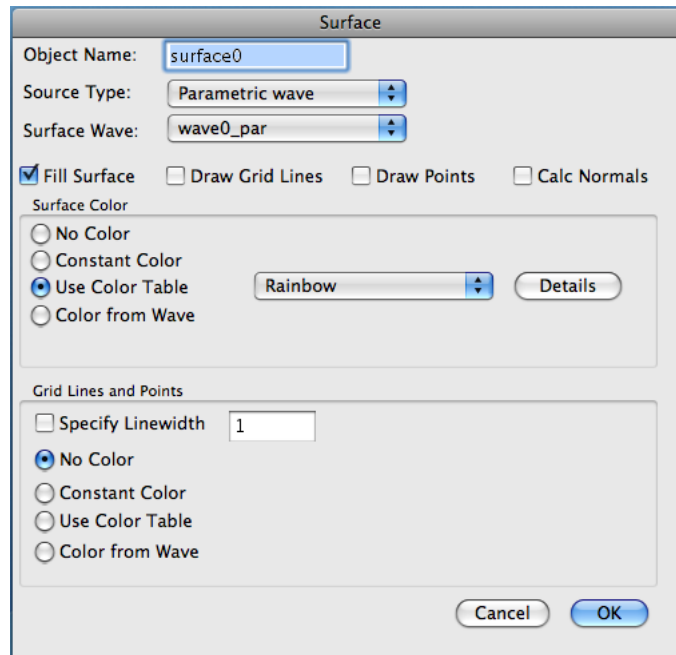
Click the “Appearance” button to generate a color image.



Use the CTRL-I (CMD-I for Mac) key sequence to manually show the cursors.

3D Surface Plots

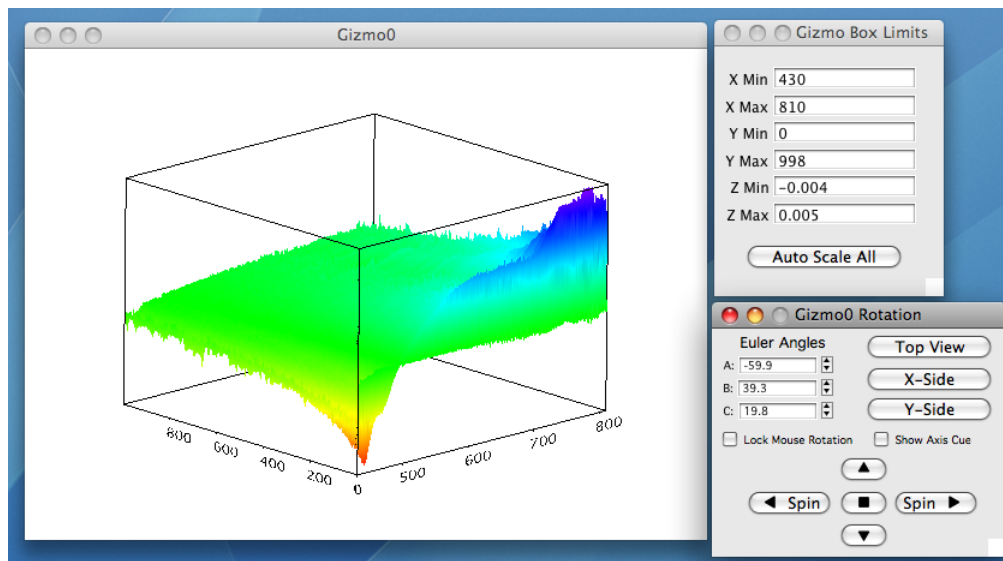
- 2) To generate a 3D Plot of your data:
Windows -> New -> 3D Plots -> Surface Plot



Modify as appropriate:

Gizmo -> Axis Range

Gizmo -> Rotation

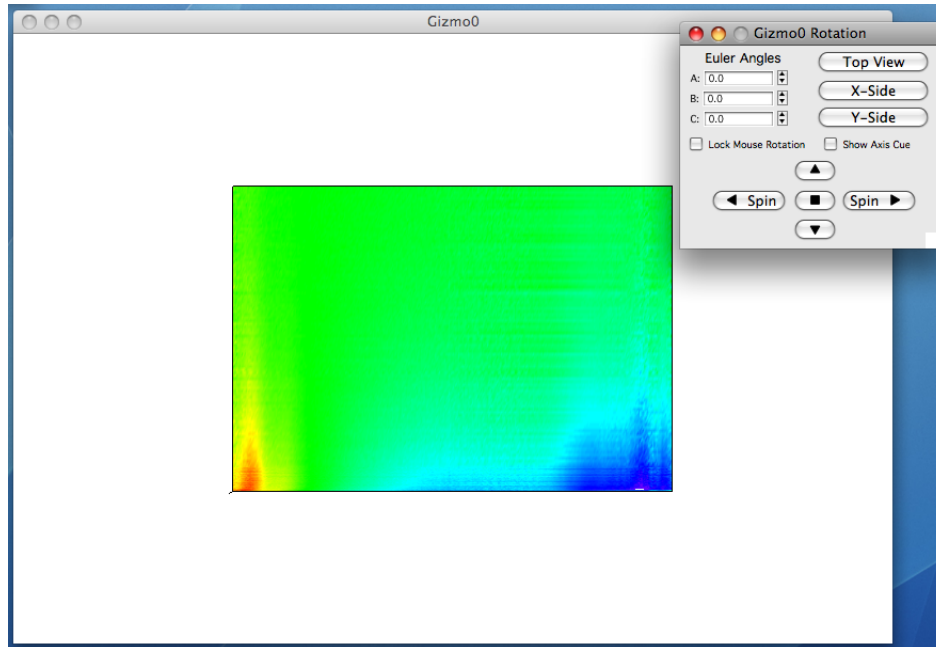


Scatter and Contour Plots

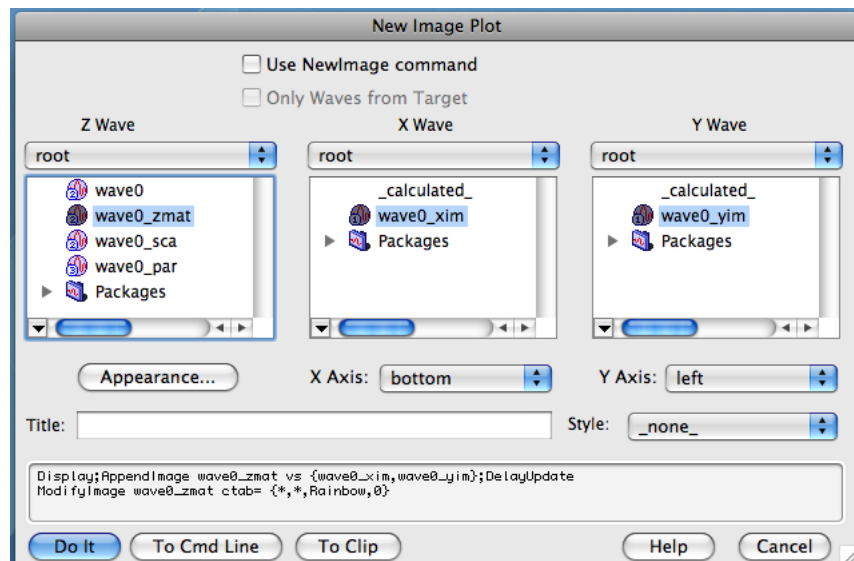
- 3) Similar procedures can be used to create 3D scatter plots and 2D contour plots.

Image Plots

- 4) Image Plots can be generated in one of the following ways:
 - A. In the Gizmo rotation panel, click "Top View."



- B. The ProcessTA procedure generates axes which are *approximately* scaled to generate an image ($N \rightarrow N+1$ points with no interpolation).
Windows \rightarrow New \rightarrow Image Plot

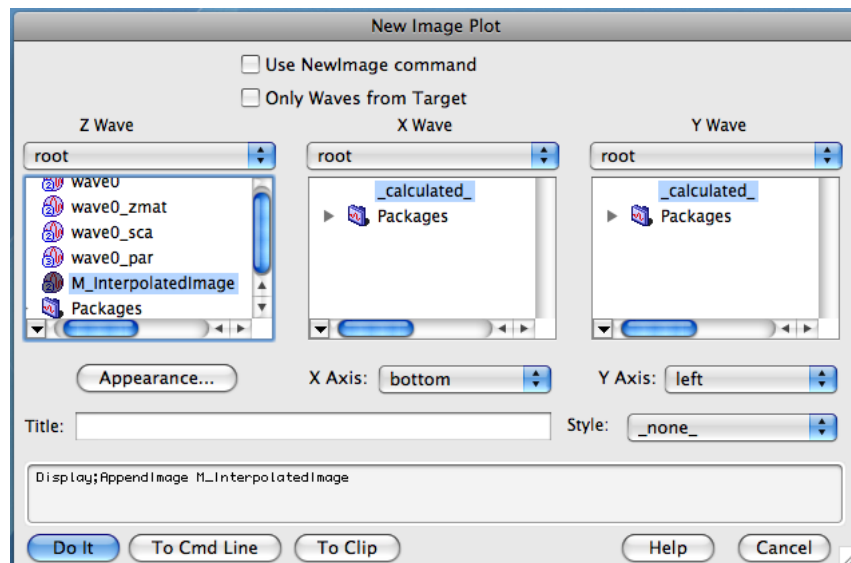


Click the “Appearance” button to generate a color image.

C. Interpolate a particular region of interest. Image interpolation can take several minutes depending on the image size and the mesh. An example command would be:

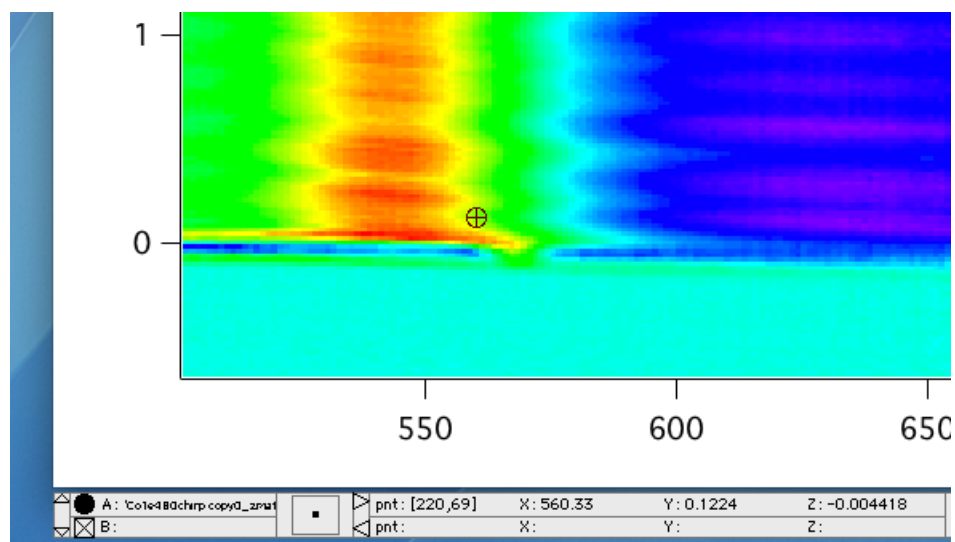
`ImageInterpolate /S={600,1,800,0,0.1,10} Voronoi wave0_sca`

This will generate an interpolated image over the spectral range 600 – 800 nm (1 nm mesh) and time range from 0 – 10 ps (100 fs mesh). On a 2.5 GHz Intel Core 2 duo processor with 4 GB of RAM, this command takes ~ 9 minutes on an original XYZ data set that contains 109056 rows.



Extracting Spectra and Kinetics from an Image Plot

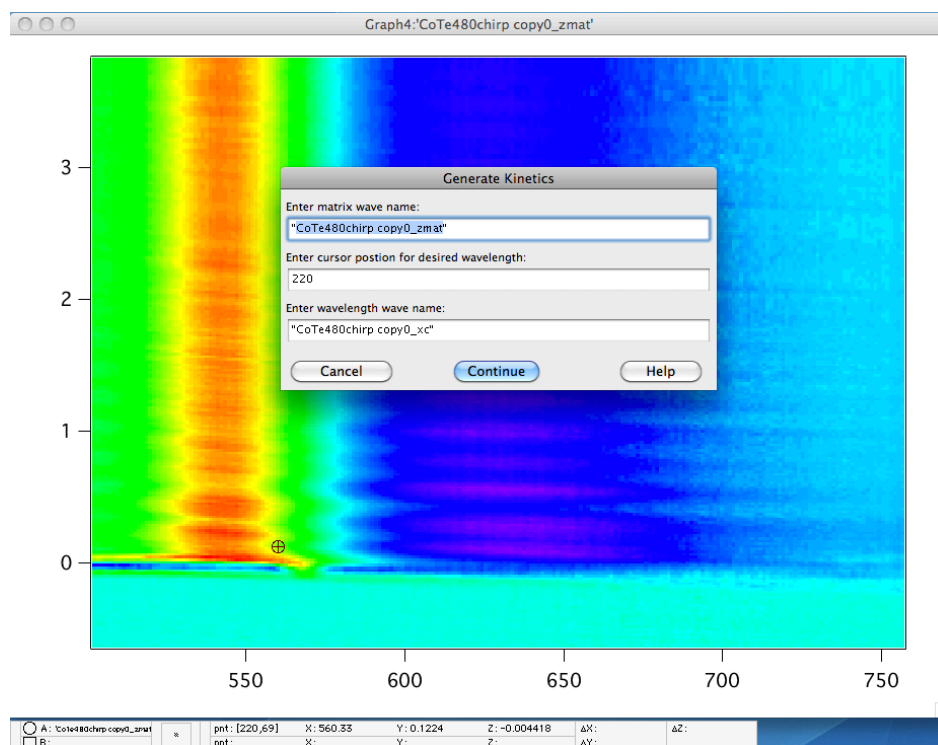
- 1) Extract Spectra and Kinetics from an image from *filename_zmat*.
The simplest way is to use the automatically generated image plot. Place the cursor in the spot where the data is to be extracted:



Select the appropriate function from the TA analysis menu:

TA Analysis → Extract Kinetics

TA Analysis → Extract Spectra



The procedure will attempt to auto-fill these values for you:

Matrix wave name: filename_zmat

Cursor position: element n [m] for (n,m) matrix for kinetics [spectra]

- note: this is not the actual wavelength or time value

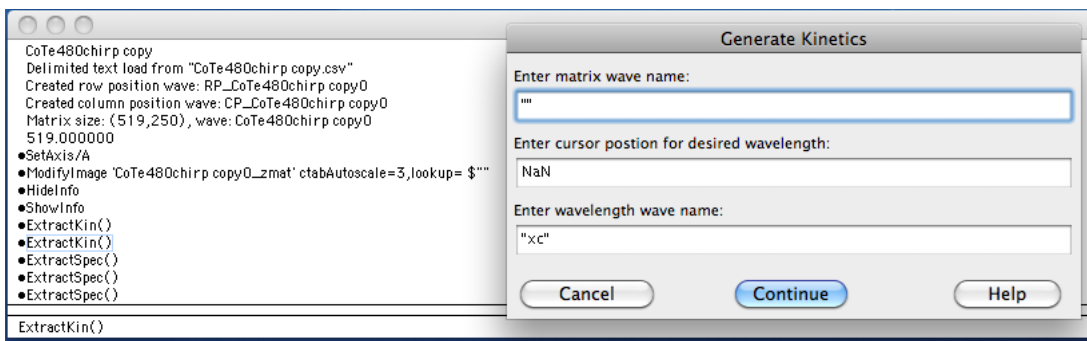
Wavelength [Time] wave name: typically *filename_xc* [*filename_yc*]

- note: this is not the waves used generate the image plot (which contain N+1 points rather than the desired N)

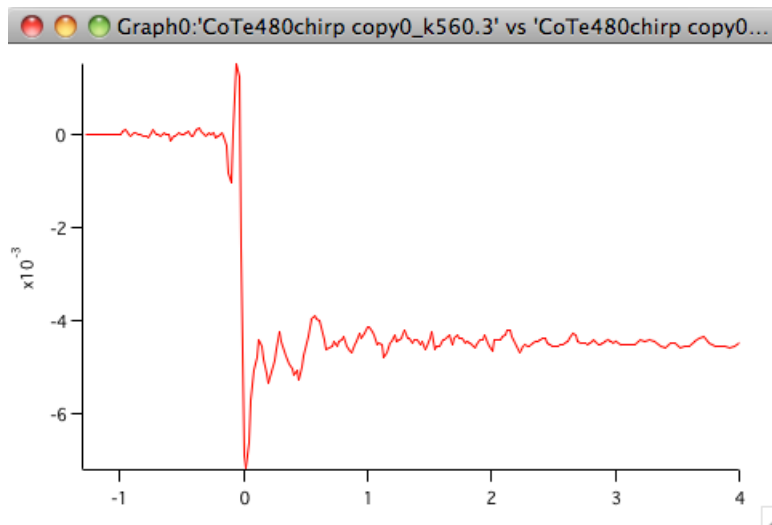
Alternatively, the functions can be called directly from the command line and the matrix information can be entered manually using the guideline above.

ExtractKin()

ExtractSpec()

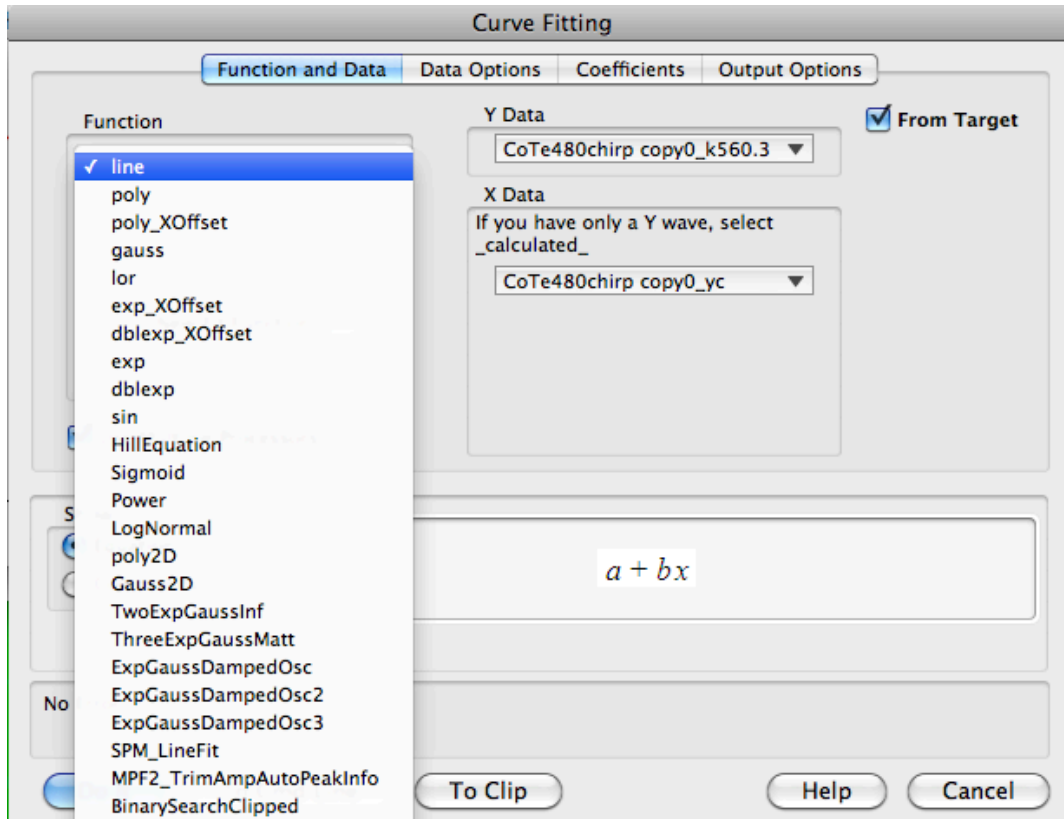


New waves will be generated with the suffix *_kwavelength* or *_stime*. 1D plots will be generated automatically if the expected file structure exists.



Curve and Peak Fitting

- 1) Fits of the kinetic data can be made using the built in Curve Fitting functions.
Analysis → Curve Fitting →



A few special fitting functions are added in this experiment:

TwoExpGaussInf
ThreeExpGaussMatt
ExpGaussDampedOsc[2,3]

These use the built-in ExpGauss function to convolute the effect of impulsive excitation and exponential decay. The only difference is that a common pulsewidth and time zero is used for all exponentials.

Parameters for ThreeExpGaussMatt are entered in the Coefficients tab:

y0: constant y offset, usually fixed at 0 during fitting

amp1: amplitude of first exponential decay

t0: time zero (time of impulsive excitation)

pw: pulsewidth can be usually set to .12 for initial guess.

- Note: This is the standard deviation of the Gaussian used to approximate the laser pulse, not the FWHM. The conversion is $\text{FWHM} \sim 2.355 \cdot \text{pw}$.

G1: the rate of the first exponential decay component

- Note: This is equal to the inverse of the lifetime

amp2: amplitude of the second exponential decay component

G2: the rate of the second exponential decay component

amp3: amplitude of the third exponential decay component

G3: the rate of the second exponential decay component

The parameters for TwoExpGaussInf are identical with amp3 now called Ainf and G3 being automatically set to 0. This is used when the last decay component is effectively flat over the time range being fit.

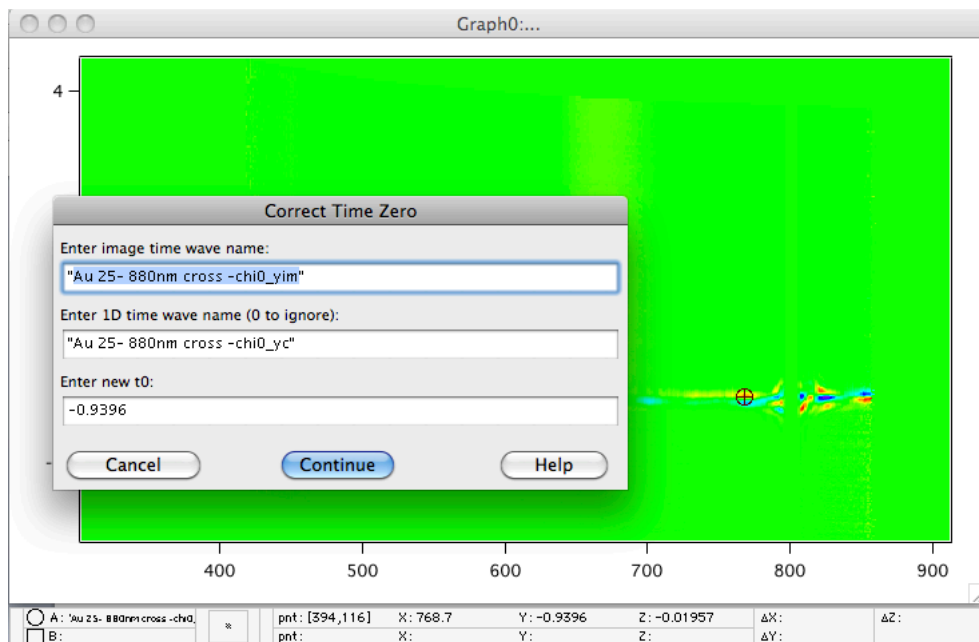
- 2) Transient spectra can be fit using the Multi-peak fit dialog.
Analysis → Multi-peak Fit → Start New Multi-peak Fit

More information can be found in the Igor Help Browser.

Data Processing Procedure

Time Zero Correction

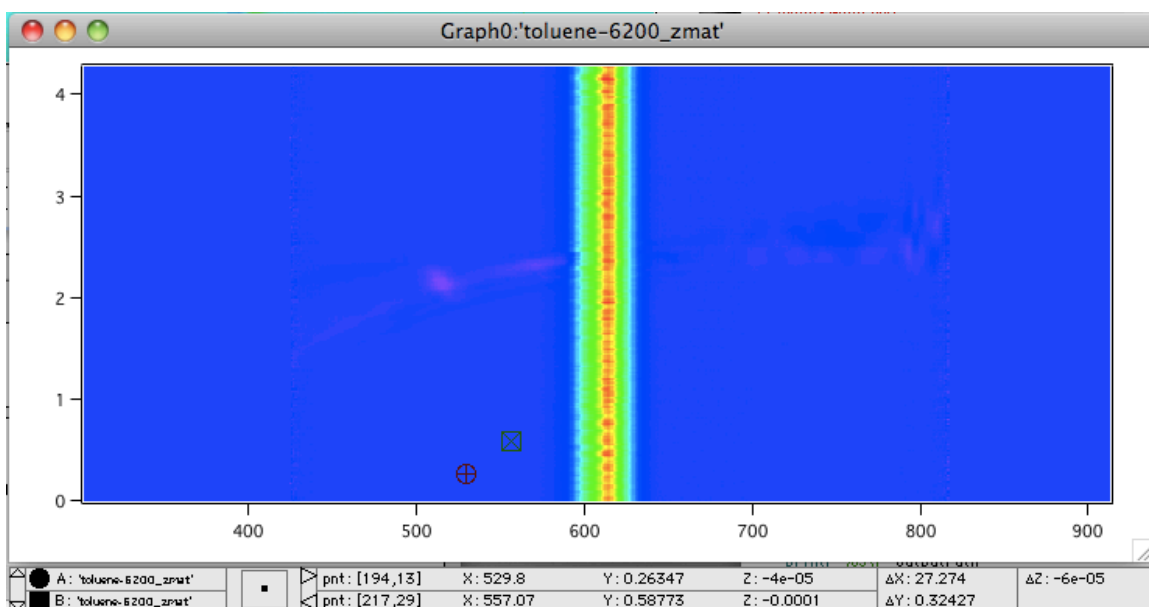
- 1) Use this procedure to correct your image and spectra for an absolute time zero. Place cursor A on your image at time zero and choose:
TA Analysis → Time Zero Correction
- 2) This procedure attempts to automatically pick the correct time files and will choose the cursor position as the new time zero. These can be modified from the dialog prompt:



- 3) The first wave name, *filename_yim* is the y axis for the image plot (N+1 point) and the second, *filename_yc*, is the y axis for extracted kinetics (N points). To leave the kinetic plots undisturbed, enter "0" in this field. To choose a different value for the new time zero, manually enter it into the appropriate field.

Scattered Light Correction

- 1) Scattered light collected by the array detectors can be automatically removed using the “Scattered Light Correction” procedure. The easiest way to do this is to use an image plot of your data.
- 2) Place cursor A and cursor B on the image to select a range to be averaged for the correction. Both of these cursors should be in the $t < 0$ regime. Avoid early times that contain empty pixels (due to chirp correction).



- 3) Choose the appropriate procedure from the TA Analysis menu:
TA Analysis → Scattered Light Correction

Get Cursor Positions

Scattered Light Correction

Cursor A: 67 Cursor B: 43

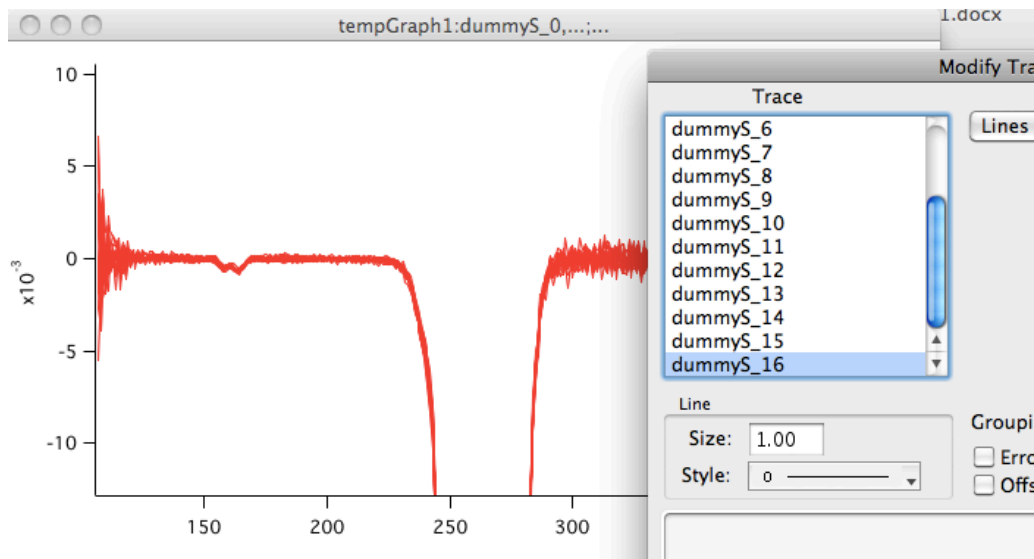
Preview Save

matrix: toluene-6200_zmat

sca_bkgd: toluene-6200_scarb

corrected: toluene-6200_zibg

- 4) Click the preview button to view the set of curves that will be averaged to make the correction file. The locations of the cursors will be sent automatically, as will the wave names.



- 5) If this range is acceptable, close the window to return to the correction dialog box and click "Save." Make sure that the image file is the top graph, i.e., click on it again to bring it to the front before clicking "Save."

Alternatively, these procedures can be run from the command line interface.

The syntax is:

PreviewSLC(filename_zmat, imin, imax) for Preview

ApplySLC(filename_zmat, imin, imax) to apply SLC

where *imin* and *imax* are the indices of the time positions to be averaged.

- 6) A new matrix is produced with the wave name *filename_zibg* along with a copy of the averaged $t < 0$ file used for the correction (*filename_scab*).

A new image plot is produced with the corrected wave file. The “Extract Spectra” and “Extract Kinetics” routines will also work with the new image file.

